

# MEASUREMENT OF THE $^{58}\text{Ni}(\text{a,g})^{62}\text{Zn}$ REACTION AND ITS ASTROPHYSICAL IMPACT

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In the innermost layers of type Ia supernovae, temperatures and densities are high enough that the material reaches nuclear statistical equilibrium. Due to the rapid expansion of the ejecta, a significant portion of this material undergoes an alpha-rich freeze-out which creates a mixture of iron-group nuclei and free alpha particles. Therefore, accurate alpha-induced reaction rates on nuclei in the iron region are important for a full understanding of the nucleosynthesis in type Ia supernovae. One such reaction, the  $^{58}\text{Ni}(\text{a,g})^{62}\text{Zn}$  reaction, was identified in a recent sensitivity study to have an impact on the final abundances of  $^{62}\text{Ni}$ ,  $^{63}\text{Cu}$ , and  $^{64}\text{Zn}$ . Cross section measurements of this reaction were performed at the University of Notre Dame using the NSCL SuN detector. The experimental results and comparison to Hauser-Feshbach calculations will be presented. The impact of using the new reaction rates in type Ia supernova nucleosynthesis models will also be discussed.